The Problem of Radiation-Reaction

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Outline

- Crash course in Electromagnetism
- Super crash course in Special Relativity
- The Problem of Radiation-Reaction
- The Lorentz-Abraham-Dirac Equation
- The Landau-Lifshitz Equation
- Study of the Landau-Lifshitz Equation

Coulomb's Law and the Electric Field

Newton's second law: $\vec{F} = m\ddot{\vec{x}}$ The force exerted by a charge Qat \vec{x}_0 on a charge q at \vec{x} is the Coulomb's force:

$$\vec{F} = K \frac{qQ}{|\vec{x} - \vec{x}_0|^3} (\vec{x} - \vec{x}_0)$$
$$:= q\vec{E}(\vec{x})$$

where K is the Coulomb's constant



Coulomb's law only works for static charges

Lorentz Force Equation

When the charge q is not at rest and has velocity \vec{v} , the force has an additional component:

$$\vec{F} = q(\vec{E} + \vec{v} \times \vec{B})$$

where \vec{E} is the electric field, and \vec{B} is called the magnetic field.



This is the Lorentz force.

Together with Newton's second law it gives the Lorentz force equation:

$$m\ddot{\vec{x}} = q(\vec{E} + \vec{v} \times \vec{B})$$

In order to fully describe a system of charges, we need to determine the electric and magnetic fields.

Marmall's Fautions



The Special Theory of Relativity (1905) <u>Postulate</u>: The speed of light c is the same <u>for all observers</u> <u>Einstein</u>: The electric and magnetic fields are the manifestation of the same field, viewed differently by different observers.

$$cdt = \sqrt{(dx)^2 + (dy)^2 + (dz)^2}$$

Introduce the notation: $x^0 = ct, x^1 = x, x^2 = y, x^3 = z$ and the matrix: $\eta_{\alpha\beta} = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{bmatrix}$



and we can rewrite the condition as: (Sum over $\eta_{\alpha\beta}dx^{\alpha}dx^{\beta} = 0$ repeated indices) Mathematically, this turns spacetime into a 4D manifold with an 'inner-product': $\langle u, v \rangle = \eta_{\alpha\beta}u^{\alpha}v^{\beta}$

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Covariant Formulation of Maxwell's equations Write the non-homogenous Maxwell's equations in components:

OT

$$\nabla \cdot \vec{E} = 4\pi\rho \implies \frac{\partial E_x}{\partial x^1} + \frac{\partial E_y}{\partial x^2} + \frac{\partial E_z}{\partial x^3} = 4\pi\rho$$

$$\nabla \times \vec{B} \xrightarrow{\partial \vec{E}} A_{\pi} \vec{r} \implies -\frac{\partial E_x}{\partial x^2} + \frac{\partial B_z}{\partial x^3} = 4\pi J_x$$
We can wr
$$\mathbf{Geometry intermission}$$
with
$$F^{\alpha\beta} = \begin{bmatrix} E_x & 0 & -B_z & B_y \\ E_y & B_z & 0 & -B_x \\ E_z & -B_y & B_x & 0 \end{bmatrix} \text{ and } J^{\alpha} = (\rho, \vec{J})$$
how do we get the homogenous equations?
four-potential:
$$A^{\alpha} = (\phi, \vec{A}) \implies F^{\alpha\beta} = \frac{\partial A^{\beta}}{\partial x_{\alpha}} - \frac{\partial A^{\alpha}}{\partial x_{\beta}}$$

Covariant Formulation of Lorentz Force equation In classical mechanics: $\vec{x} = (x(t), y(t), z(t))$ In special relativity: $x^{\alpha} = (x^{0}(\tau), x^{1}(\tau), x^{2}(\tau), x^{3}(\tau))$ The parameter τ is the proper time, the time measured in the reference frame of the particle itself. IME The generalization of the velocity is UTURE LIGHT COM the <u>four-velocity</u>: $u^{\alpha} = \frac{dx^{\alpha}}{d\tau} = \gamma(1, \vec{v})$ OBSERVER ERSURFACE OF where $\gamma = \frac{1}{\sqrt{1-v^2}}$ PAST LIGHT CONE The Lorentz force equation $m\dot{\vec{v}} = q(\vec{E} + \vec{v} \times \vec{B})$ turns into: $m\dot{u}^{\alpha} = qF^{\alpha\beta}u_{\beta}$

The Problem of Radiation-Reaction

The Lorentz force equation: $m\dot{u}^{\alpha} = -eF^{\alpha\beta}u_{\beta}$

Maxwell equations => the rate at which energy is radiated away from the electron is:

$$\mathcal{R} = -m\tau_0 \dot{u}^{\alpha} \dot{u}_{\alpha} \qquad \qquad \tau_0 = \frac{2}{3} \frac{e^2}{m} = 6.24 \times 10^{-24} \,\mathrm{s}$$

Therefore an accelerating charge loses energy.

This effect is not included in the Lorentz force equation!

<u>Example</u>: circularly polarized plane wave





Lorentz force equation doesn't account for energy lost

The Lorentz-Abraham-Dirac (LAD) Equation <u>Dirac (1938)</u>: Maxwell equations and energy conservation give

$$m\dot{u}^{\alpha} = -eF^{\alpha\beta}u_{\beta} + m\tau_{0}\left[\ddot{u}^{\alpha} + \dot{u}^{2}u^{\alpha}\right]$$
$$F_{Lorentz} F_{RR}$$

This is the LAD equation (Lorentz-Abraham-Dirac). The 3rd order time derivative requires another initial condition (initial acceleration), and results in infinitely many non-physical solutions...

Dirac replaced the additional condition with an "asymptotic condition". Instead of giving the initial acceleration, give the <u>final acceleration</u>.

The Landau-Lifshitz (LL) Equation The LAD: $m\dot{u}^{\alpha} = -eF^{\alpha\beta}u_{\beta} + m\tau_0 \left[\ddot{u}^{\alpha} + \dot{u}^2 u^{\alpha}\right]$ If τ_c is a characteristic time scale, the radiation-reaction force is of order $\varepsilon = \frac{70}{-2}$ So to leading order in ε , we get the Lorentz force Eq: $m\dot{u}^{\alpha} = -eF^{\alpha\beta}u_{\beta} + O(\varepsilon)$ Therefore,

$$m\ddot{u}^{\alpha} = -e\frac{d}{d\tau}\left(F^{\alpha\beta}u_{\beta}\right) + O(\varepsilon)$$

Using this approximation in the LAD gives the Landau-Lifshitz equation:

 $m\dot{u}^{\alpha} = -eF^{\alpha\beta}u_{\beta} - e\tau_0 \left\{ F^{\alpha\beta}_{,\gamma}u_{\beta}u^{\gamma} - e/m \left[F^{\alpha\beta}F_{\beta\gamma}u^{\gamma} - F^{\beta\gamma}F_{\gamma\delta}u^{\delta}u_{\beta}u^{\alpha} \right] \right\}$

History of the Radiation-Reaction Problem

- 1865 : Maxwell formulates the theory of electromagnetism
- 1892 : Lorentz introduces the Lorentz force equation and argues that the radiation-reaction force is missing...
- 1903 : Abraham finds the radiation-reaction force
- 1904: Lorentz assumes that the electron is a solid sphere, and proves that the self-force exerted by the sphere on itself = radiation-reaction force
- 1905: Poincaré realizes that this means that in the lack of other forces, the sphere will explode...
- 1905-1920: People realize that the atom will be unstable
- 1920's: Schrödinger and his buddies formulate Quantum Mechanics. For about a decade people "forgot" about radiation-reaction...
- 1938: Dirac derives the covariant Lorentz-Abraham-Dirac equation
- 1938-1948: The physics community notices that the LAD has non-physical solutions
- 1948: Eliezer "derives" a new equation
- 1951: Landau & Lifshitz introduce the LL equation
- 1956: Caldirola claims that time is discrete, and replaces the ODE with a finite-difference equation making all the numerics people happy...
- 1962: Prigogine & Henin hypothesize a new equation









History of the Radiation-Reaction Problem

- 1964: Nodvik hypothesizes a new equation
- 1970: Teitelboim hypothesizes a new equation
- 1971: Mo & Papas modify energy-momentum conservation and hypothesize a new equation
- 1972: Leiter criticizes Mo & Papas. "One cannot modify the law of energy conservation..."
- 1973: Mo & Papas to Leiter "Only experiment will tell what is energy conservation"
- 1976: Gonzales & Gascon claim that LAD is only an approximated equation and derive a new equation.
- 1977: Petzold & Sorg generalize Caldirola's equation
- 1981: Valle et al. claim that Mo-Papa's equation is the correct equation
- 1988: Valentini proves the non-physical solutions of LAD are due to non-analytic fields
- 1992: Yaghjian derives an equation for a spherical particle
- 1997-2000: Rohrlich claims that Yaghjian's equation is the correct equation
- 2008: Rohrlich "Using physical arguments, I derive the physically correct equations of motion for a classical charged particle from the LAD equation which is well known to be physically incorrect."
- 2009: Gralla, Harte & Wald rederive Landau-Lifshitz equation <u>rigorously</u> using distribution theory
- 2009: Sokolov et al. introduce another equation

Mathematical Origins of the Problem of Radiation-Reaction

Maxwell's equations: $\partial_{\alpha}F^{\alpha\beta} = 4\pi J^{\beta}$ Lorentz force equation: $m\dot{u}^{\alpha} = -eF^{\alpha\beta}u_{\beta}$

There are three different kind of problems: (1) Studying the evolution of the electromagnetic fields for known sources J^{α} : it has a well-posed initial value formulation.

Also for a point-particle: $J^{\alpha}(x) = -eu^{\alpha}\delta(x-z(t))\frac{d\tau}{dt}$

(2) Study the motion of a particle in a given external electromagnetic field $F^{\alpha\beta}$: it also has a well-posed initial value formulation.

(3) The coupled system: doesn't make mathematical sense because the field is necessarily singular at the position of the particle The Radiation-Reaction Dominated Regime Solve the Landau-Lifshitz equation <u>analytically</u> for a laser wave.

Recall that the rate at which energy is radiated is: $\mathcal{R} = -m\tau_0 \dot{u}^{\alpha} \dot{u}_{\alpha}$

for the solution of Landau-Lifshitz:

 $\mathcal{R} = -\frac{2}{3}e^{2}\frac{(k \cdot u)^{4}}{(k \cdot u_{0})^{2}} \left\{ a_{0}^{2}\hat{A}^{\prime 2} \quad \text{(Lorentz)} \right. \\ \left. +\tau_{0}(k \cdot u_{0}) \left[2a_{0}^{2}\hat{A}^{\prime \prime} \cdot \hat{A}^{\prime} - 2a_{0}^{4}\psi \hat{A}^{\prime 2} \right] \right\}$

Landau-Lifshitz correction ~ $\omega \tau_0 a_0^4$ For a typical laser $\omega \tau_0 \sim 10^{-8}$ $a_0 = \frac{eA}{m} \text{ intensity of laser}$ $\hat{A}^{\alpha} \text{ normalized 4-potential}$ $k^{\alpha} \text{ wave 4-vector}$ $u_0^{\alpha} \text{ initial 4-velocity}$

So Radiation-reaction effects are important when: $a_0^2 \sim 10^8$ This is <u>far</u> beyond current technological capabilities $a_0 \sim 10$

The Radiation-Reaction Dominated Regime

$$u^{\alpha} = \gamma(1, \vec{v}) \Longrightarrow u^{\alpha} = O(\gamma)$$

This scales radiation-reaction effects by a factor γ_0 !



RRDR Criterion: $\gamma_0 a_0^2 \sim 10^8$

$$\Delta = \frac{1}{2\pi} \int_0^{2\pi} \frac{|E_{LL}(\xi) - E_{Lorentz}(\xi)|}{|E_{LL}(\xi) + E_{Lorentz}(\xi)|}$$



Collaboration

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- Lance Labun (Physics department UoA)
- N. Elkina, C. Klier & H. Ruhl (LMU Munich)

Thank you!

Pictures for Linearly Polarized Wave





The Setup



We solve the LL eq. for the potential: $A^{\alpha}(x) = A_0 \operatorname{Re}\left[\varepsilon^{\alpha} f(\xi)\right]$

with:

- Polarization vector ε^{α} Wave vector k^{α} $\xi = k \cdot x$ k^{α} $k^{2} = 0$ $|\varepsilon|^{2} = -1$ $k \cdot \varepsilon = 0$

Examples Linear Polarization $\varepsilon^{\alpha} = (0, 1, 0, 0)$ $k^{\alpha} = (\omega, 0, 0, k)$ $f(\xi) = \sin(\xi)$ $\vec{A} = -A_0 \sin(kz - \omega t)\hat{x}$

Circular Polarization $\varepsilon^{\alpha} = \frac{1}{\sqrt{2}}(0, 1, -i, 0)$ $k^{\alpha} = (\omega, 0, 0, k)$ $f(\xi) = \sqrt{2}e^{i(\xi-\xi_0)}$ $\vec{A} = A_0 \left[\cos(kz - \omega t) \hat{x} \right]$ $-\sin(kz-\omega t)\hat{y}$



Solution of the Landau-Lifshitz Equation The LL equation is non-linear and therefore very difficult to solve in terms of proper time... $m\dot{u}^{\alpha} = -eF^{\alpha\beta}u_{\beta} - e\tau_0 \{F^{\alpha\beta}_{,\gamma}u_{\beta}u^{\gamma} - e/m [F^{\alpha\beta}F_{\beta\gamma}u^{\gamma} - F^{\beta\gamma}F_{\gamma\delta}u^{\delta}u_{\beta}u^{\alpha}]\}$ "Trick": change variables $\tau \mapsto \xi = k \cdot x$ $\frac{d\xi}{d\tau} = k \cdot u$

 $(k \cdot u) u^{\prime \alpha} = -\frac{e}{m} F^{\alpha \beta} u_{\beta} - \frac{e}{m} \tau_0 \left\{ F^{\alpha \beta}_{,\gamma} u_{\beta} u^{\gamma} - \frac{e}{m} \left[F^{\alpha \beta} F_{\beta \gamma} u^{\gamma} - u_{\beta} F^{\beta \gamma} F_{\gamma \delta} u^{\delta} u^{\alpha} \right] \right\}$

SOLUTION